# The national X-ray absorption spectrometers FAME and FAME-UHD after the ESRF-EBS phase



French Absorption spectroscopy beamline for Materials and Environmental sciences



Ultra High Dilution







# F\_CRG pool



### FAMEs group



### Hard X-ray Absorption Spectroscopy



### FAME / FAME-UHD: common optical layout



### FAME / FAME-UHD: detection schemes

X-ray Fluorescence X-ray Emission Spectroscopy 16 elastic (d) - SSD (e) (f) CAS Ka, & Ka, Co K Co K Co K. Co K CAS KB (x10) unit.) 1.2 Emission intensity (arb. unit.) inelasti (arb. **HERFD-XAS EXAFS** 0.8 +XES Co K 0.6 **XANES** Co K 0.4 +RIXS 0.2 7500 8000 8500 5500 6500 7000 6940 7650 7660 7670 6900 6930 6950 7630 7640 E<sub>emission</sub> (eV) E<sub>emission</sub> (eV) E<sub>emission</sub> (eV) rone CAS لالم Detector Energy dispersion with a solid  $\Delta E_{\text{emission}} \approx 1-3 \text{eV}$ state detector (electronic)  $\Delta E_{\text{emission}} \approx 150-300 \text{eV}$ Wavelength dispersion with a sample SSD or SDD sample crystal analyser spectrometer natic boom (optical) ΔE<sub>incident</sub> ≈ 0.5-1eV ΔE<sub>incident</sub> ≈ 0.5-1eV

### FAME / FAME-UHD: sample environment



# FAME: state-of-the art EXAFS

Environmental Science & Biology: Ag-NPs in sewage sludge applied on agricultural soils



- XANES (PCA + LCF) and EXAFS
- complementary to XRF/TEM microscopy



Ana E. Pradas del Real et al., "Fate of Ag-NPs in Sewage Sludge after Application on Agricultural Soils", *Environ. Sci. Technol.* 2016 **50**(4), 1759-1768

### Detection limit: 10-100 ppm

Hazemann et al., "High Resolution Spectroscopy on an X-ray Absorption Beamline", J. Synchr. Rad. 16 (2009) 283-292 Mauro ROVEZZI (mauro.rovezzi@esrf.fr), CNRS-OSUG | 2017-11-28 | #7

# FAME-UHD: lower the detection limit + HERFD-XAS



### Detection limit: ≈0.5 ppm

Proux et al., "High-Energy Resolution Fluorescence Detected X-ray Absorption Spectroscopy: A Powerful New Structural Tool in Environmental Biogeochemistry Sciences", J. Env. Quality **46** (2017) 1146-1157

# EBS for a bending magnet spectroscopy beamline

### Short bending magnet source

Peak field: 0.856 T Front end aperture: 2.2 mrad



#### Ray tracing simulations (Shadow)









# 2019: heavy works on the beamlines!



### **FAME**

- FIP migrates on BM07  $\rightarrow$  new spaces
  - Beamline moves up to 146 mm

FAME-UHD

Beamline moves up to 40 mm

Optics ready to take the power of the new source  $\rightarrow$  restart in 2020

# FAME after ESRF-EBS

Micro-spectroscopy at 1×1  $\mu m^2$ 

- New optics: mirrors and sagittal bender
  - Full flux in 30×30  $\mu$ m<sup>2</sup>
  - Secondary source + KB mirrors
- Expected gain in flux density up to ≈60



• Compact sample environments (e.g. diamond anvil cell)



### FAME-UHD after ESRF-EBS

### Gain in flux density

- Full flux in 30×30  $\mu m^2$
- Better energy resolution
- Lower background (scattering)



### Spectrometer with 14 spherically bent crystal analyzers

- Commissioning in January 2018
- Procurement of new full sets of crystal analyzers

# Conclusions

ESRF-EBS		EXAFS / XANES	HERFD-XAS
High throughput spectroscopy on natural samples	High Dilution	10 ppm	0.5 ppm and below
	Inhomogeneity	≈30×30 µm² (full flux) ≈1×1 µm2 (µbeam)	≈30×30 μm² (full flux)

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Next call for proposals: March 2018

Thank you for your attention

### Choice of the new source: SBM, 3PW or 2PW?

**2PW** 

SBM







**3PW** 





≈40% out of focus

# Monochromator sagittal bender and crystal design



### Users and scientific fields

